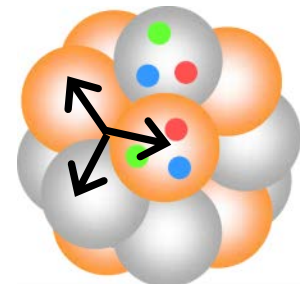
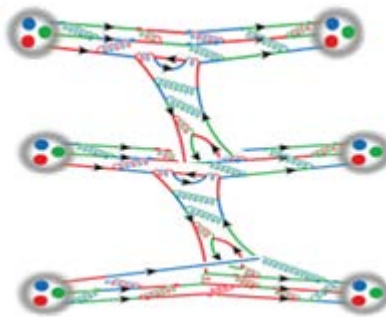
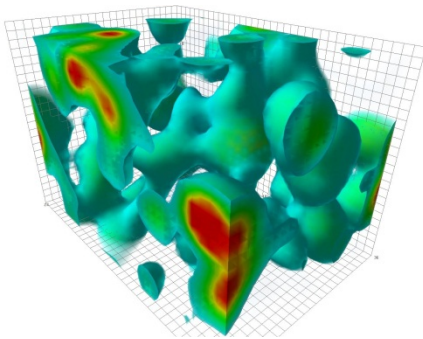


Nucleon-Nucleon Interactions from Lattice QCD

Takumi Doi
(Nishina Center, RIKEN)



Outline of the Lecture

- Lecture 1

- Introduction
- Review of lattice QCD simulations (c.f. R.Gavai)
- Quick overview of the framework (HAL QCD method)
- Review of scattering problems

- Lecture 2 (tutorial)

- Nambu-Bethe-Salpeter (NBS) wave function
 - Derivation of it's asymptotic behavior
- Scatterings on the lattice
 - Derivation of Lushcer's formula

Outline of the Lecture

- Lecture 3

- Application to Nucleon-Nucleon (NN) interaction
 - Energy independent potential
- Other two-baryon interactions w/ hyperons

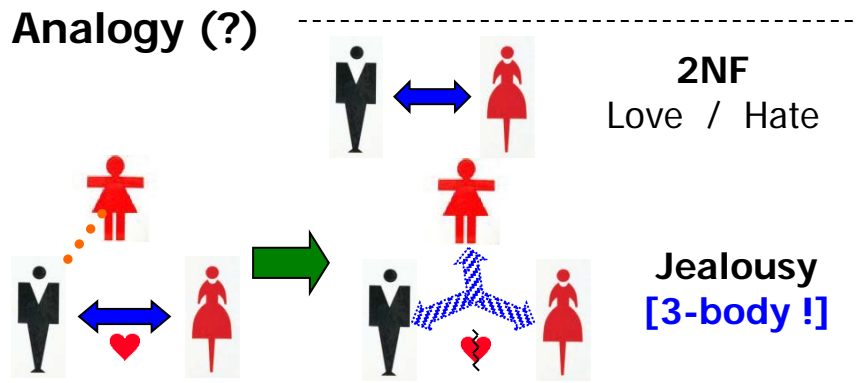
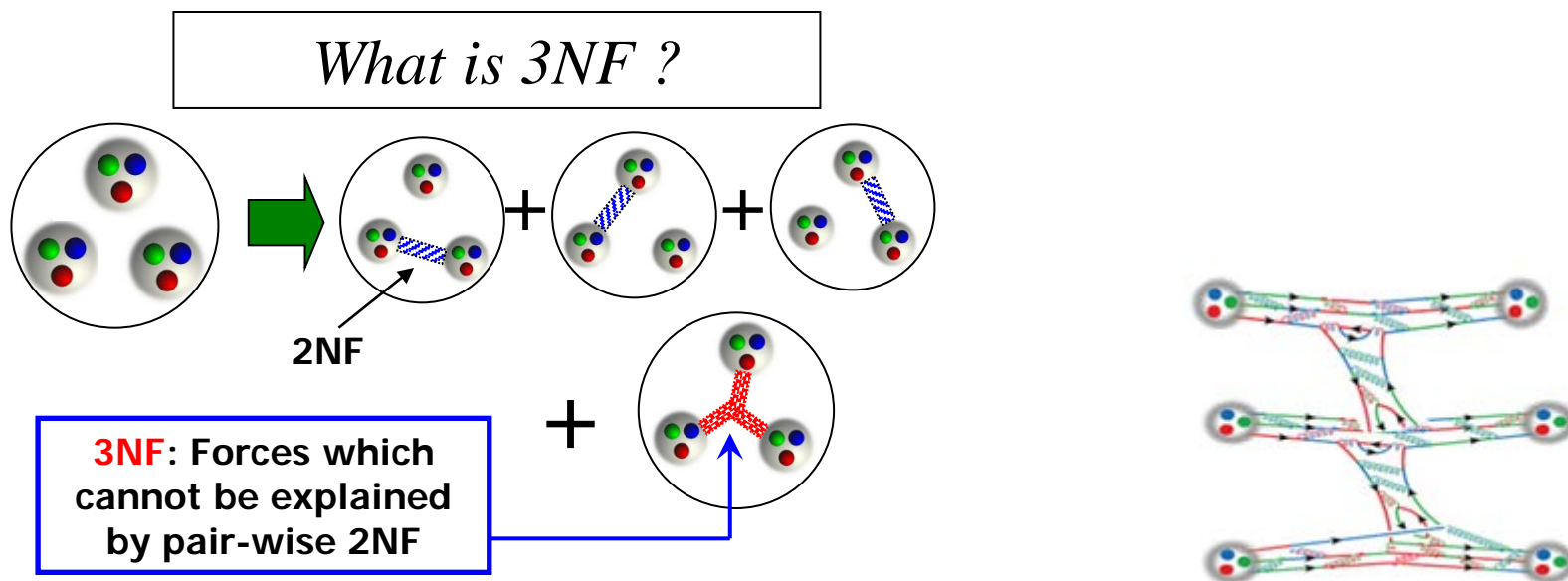
- Lecture 4

- Application to Three-Nucleon (3N) interaction
 - Unified contraction algorithm
- Summary / Prospects

Any questions are welcome !

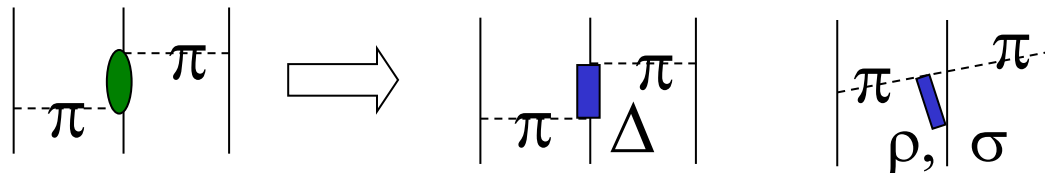
Frontier in Hadron-Hadron Interactions

⇒ **Three-Nucleon Forces (3NF)**



Three-Nucleon Forces (3NF)

- It is natural to expect the existence of 3NF
- It is very nontrivial to determine 3NF from QCD
- $2\pi E$ -3NF Fujita-Miyazawa, PTP17(1957)360
 - Off-energy-shell πN scatt



- Phenomenological models
 - Fujita-Miyazawa, Tucson-Melbourne, Urbana/Illinois, ...
- EFT expansion \rightarrow 3NF appear at NNLO
- Are 3NF really important in physics ?

	2N forces	3N forces	4N forces
LO		—	—
NLO			—
N ² LO			—
N ³ LO			
	+ ...	+ ...	+ ...

U.v.Kolck, PRC49(1994)2932

Epelbaum, Prog.Part.Nucl.Phys.57(06)654

Frontier in Hadron-Hadron Interactions

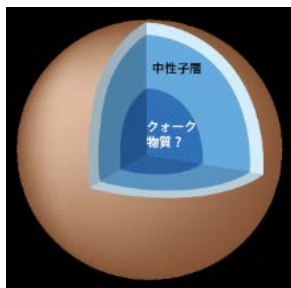
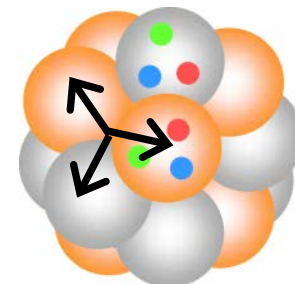
⇒ Three-Nucleon Forces (3NF)

◆ *B.E. of light nuclei*

◆ *Neutron rich nuclei*

→ *Nucleosynthesis*

◆ *Nucleon-deuteron scattering*



◆ *EoS of nuclear matter*

→ *Saturation point*

→ *Neutron Star / SuperNova*

3NF in Few-Body Systems

- Precise few-body calc:
 - e.g. benchmark calc of ${}^4\text{He}$ by 7 methods (2NF only)

Method	$\langle T \rangle$	$\langle V \rangle$	E_b	$\sqrt{\langle r^2 \rangle}$
FY	102.39(5)	-128.33(10)	-25.94(5)	1.485(3)
CRCGV	102.30	-128.20	-25.90	1.482
SVM	102.35	-128.27	-25.92	1.486
HH	102.44	-128.34	-25.90(1)	1.483
GFMC	102.3(1.0)	-128.25(1.0)	-25.93(2)	1.490(5)
NCSM	103.35	-129.45	-25.80(20)	1.485
EIHH	100.8(9)	-126.7(9)	-25.944(10)	1.486

➔ 0.5% prec. for B.E.

H.Kamada et al.,
PRC64(2001)044001

- 2N force cannot reproduce B.E.

$\delta\text{B.E.} = 0.5\text{-}1\text{MeV}$ for ${}^3\text{H}$

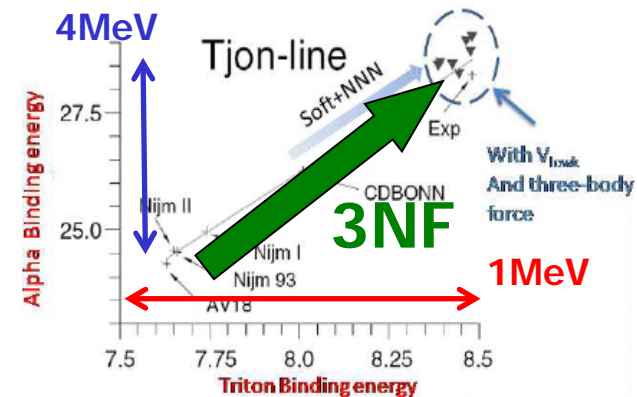
$\delta\text{B.E.} = 2\text{-}4$ MeV for ${}^4\text{He}$

missing

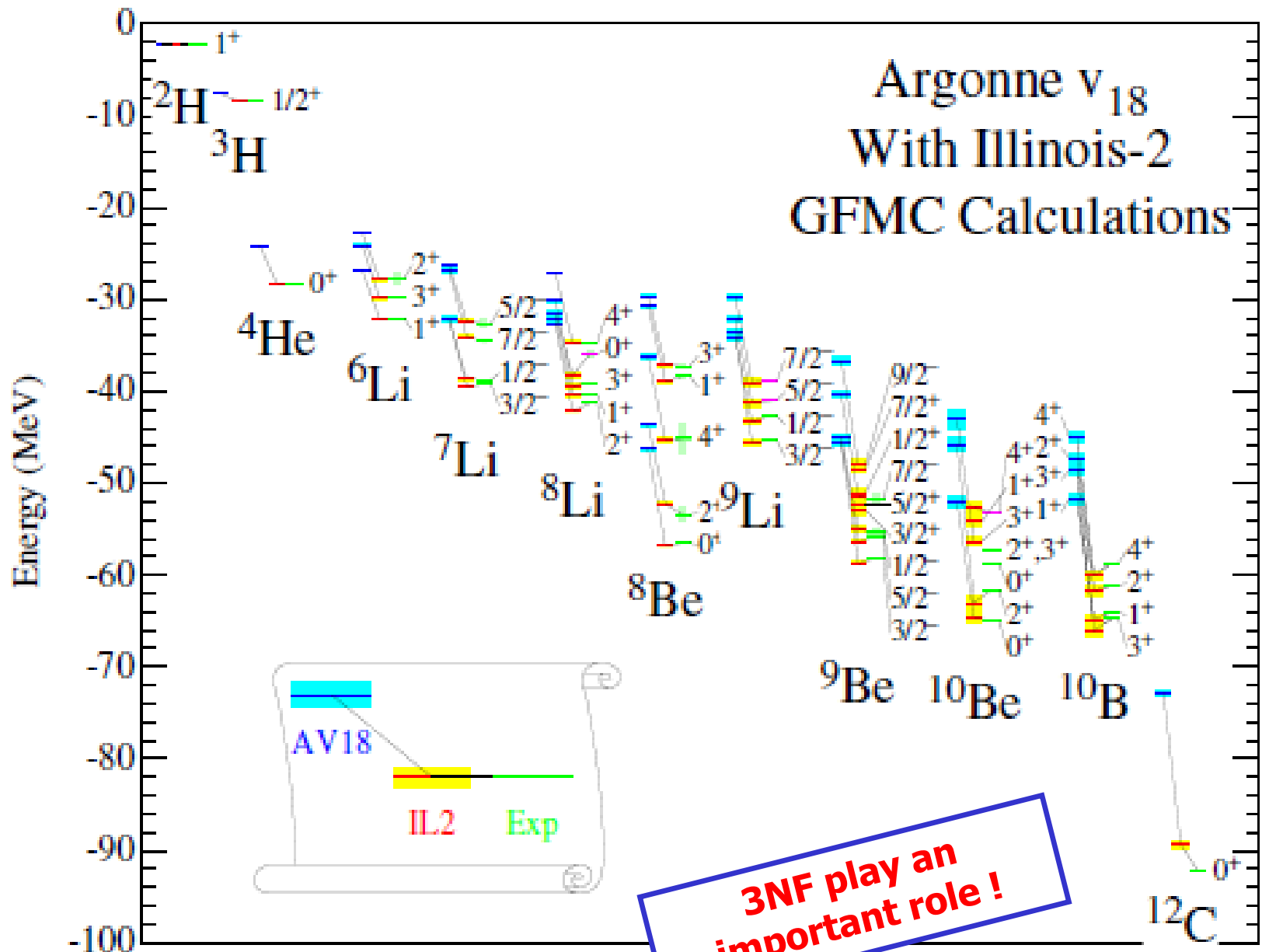
c.f. $\delta\text{B.E.} = 0.35\text{MeV}$ for ${}^3\text{H}$ in fss2



Attractive 3NF
necessary



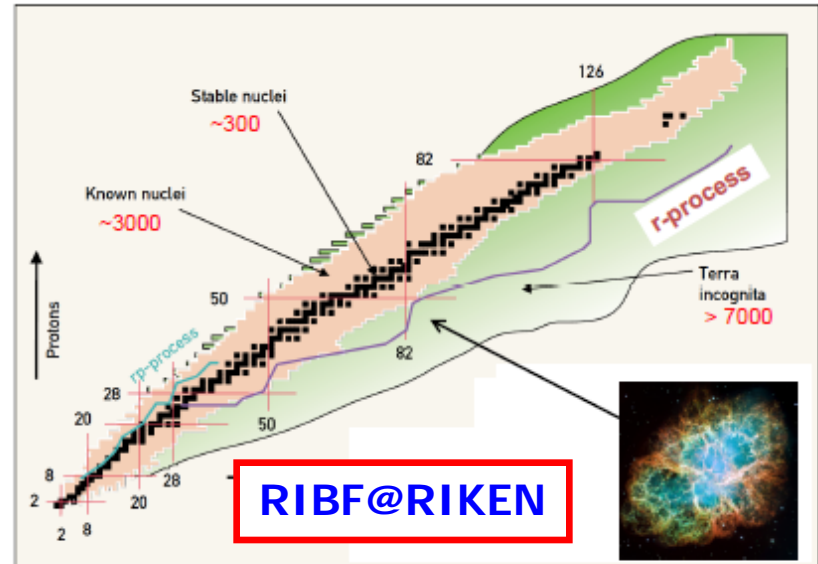
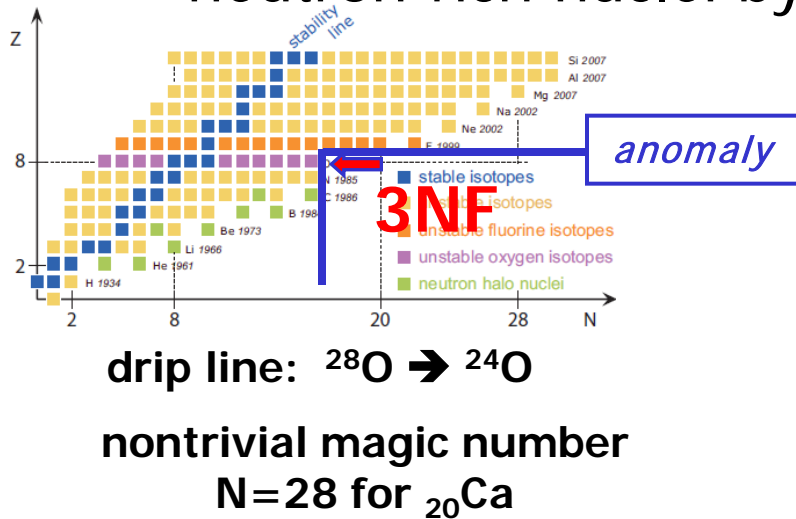
Nogga et al., PRL85(2000)944



S.C.Pieper, Riv.Nuovo.Cim31(2008)709
arXiv:0711.1500

3NF in Neutron-Rich Nuclei

- The effect on the nuclear chart
 - Anomaly in drip line and nontrivial magic number in neutron-rich nuclei by **3NF**

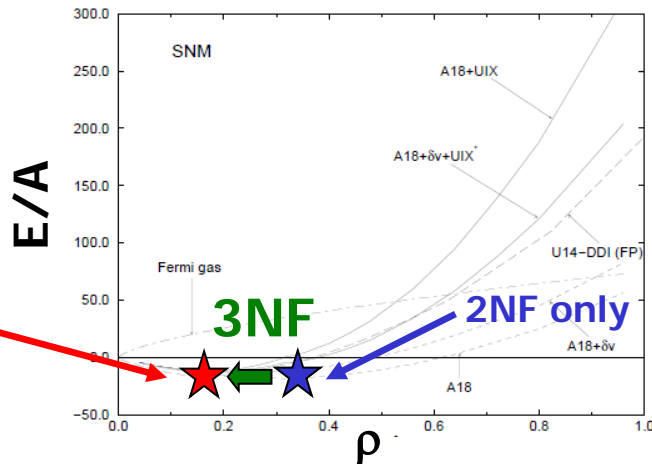


T.Otsuka et al., PRL105(2010)032501
 J.D.Holt et al., arXiv:1009.5984

Nucleosynthesis by Supernova

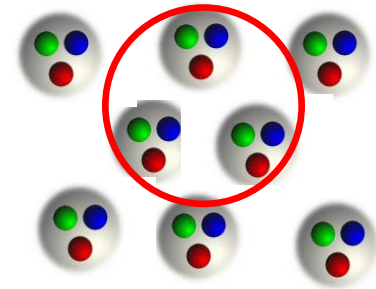
3NF in symmetric nuclear matter

- Saturation point of nuclear matter requires 3NF



Repulsive 3NF
also necessary

A.Akmal et al., PRC58(1998)1804



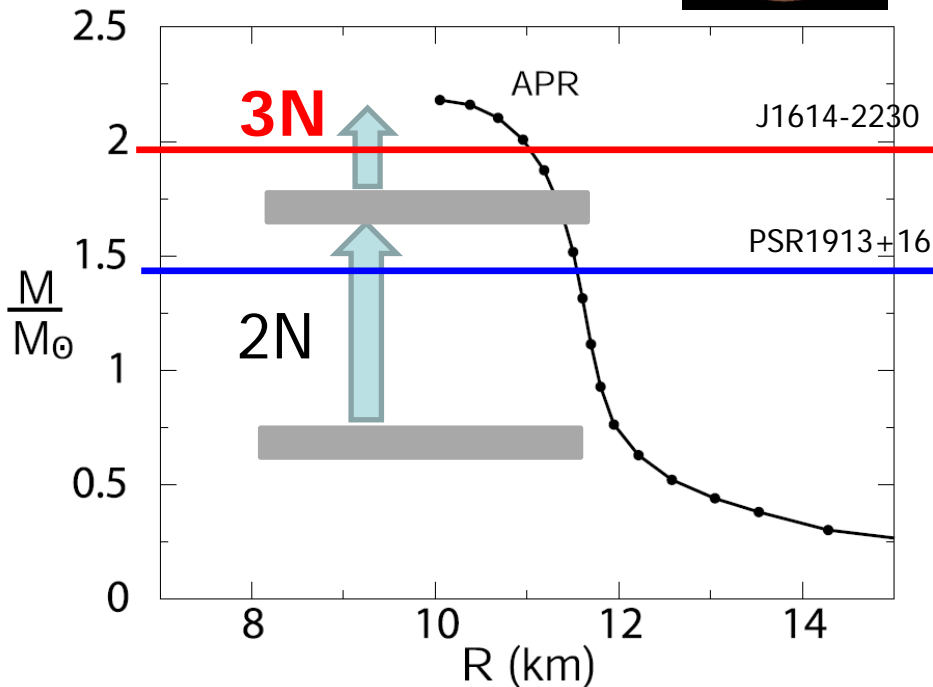
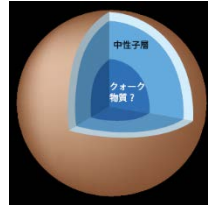
$$\rho_0 = 0.16 \text{ fm}^{-3}$$

$$E/A = 16 \text{ MeV}$$

Crucial role of 3NF in nuclear matter

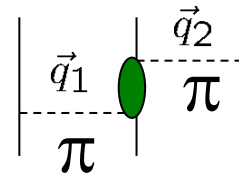
Neutron Star

(Densest system in the Universe)



Short-range repulsive 3NF is required

In phenomenological models:



2π E-3NF:

→ too strong attraction

Tucson-Melbourne

$$\sim \frac{F_\pi^2(\vec{q}_1^2)(\vec{\sigma} \cdot \vec{q}_1)}{\vec{q}_1^2 + m_\pi^2} F(\vec{q}_1, \vec{q}_2) \frac{F_\pi^2(\vec{q}_2^2)(\vec{\sigma} \cdot \vec{q}_2)}{\vec{q}_2^2 + m_\pi^2}$$

$$F_\pi(\vec{q}^2) = \left(\frac{\Lambda_\pi^2 - m_\pi^2}{\Lambda_\pi^2 + \vec{q}^2} \right) \quad \Lambda_\pi \sim 700 \text{ MeV} \ll 1.3 \text{ GeV}$$

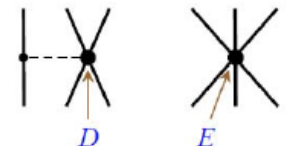
→ (effective) short-range repulsion by cut-off

Urbana/Illinois

2π E-3NF + explicit short-range repulsive term

Chiral EFT

Short-range LECs fitted



Can we understand it directly from QCD ?

EoS of Neutron Stars through Gravitational Waves ?

- Full GR simulation of binary neutron star mergers w/ Shen-EoS (stiff) and Hyperon(Λ)–EoS (soft)

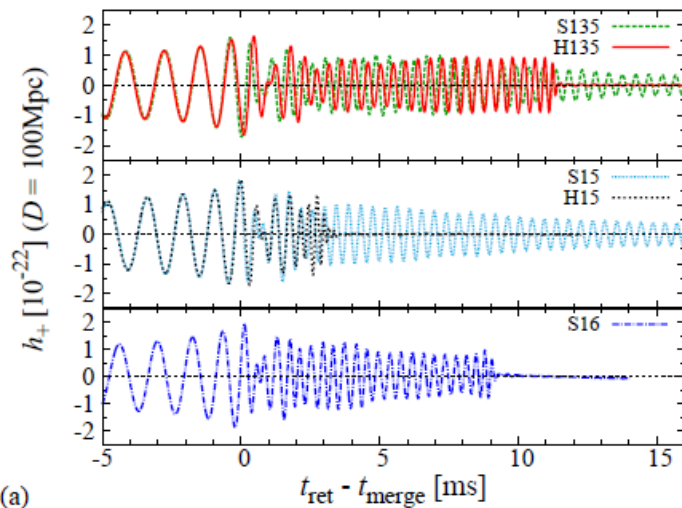


FIG. 4: (a) GWs observed along the axis perpendicular to t , $D = 100$ Mpc. (b) The effective amplitude of GWs defined by noise amplitudes of a broadband configuration of Advanced LIGO and Large-scale Cryogenic Gravitational wave Telescope (LCGT).

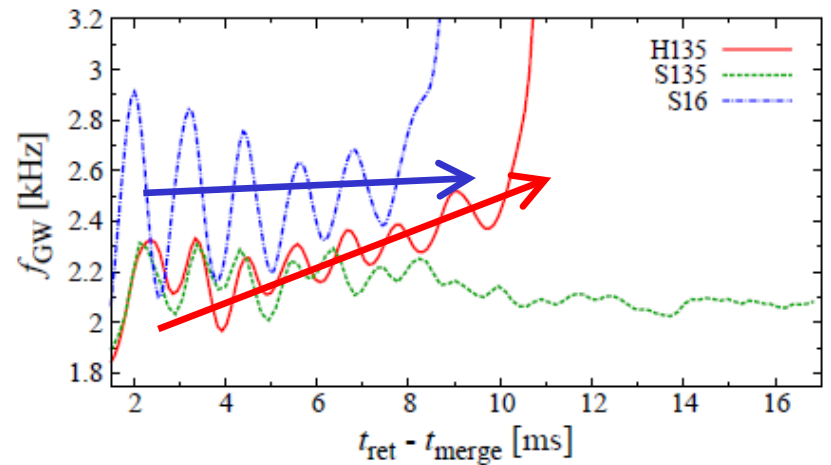
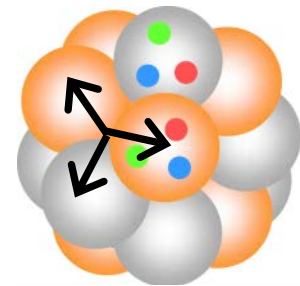
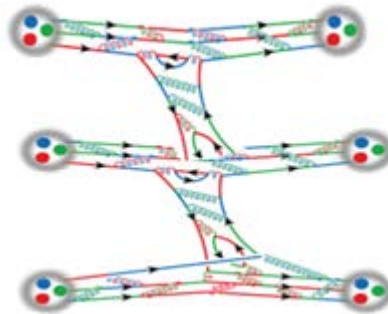
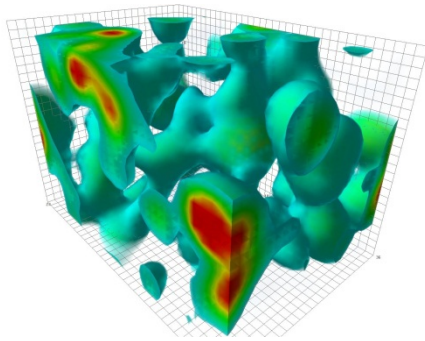


FIG. 5: $f_{GW}(t)$ in the HMNS phase, smoothed by a weighted spline, for H135 (solid red), S135 (dashed green), and S16 (dashed-dotted blue).

Y.Sekiguchi et al., arXiv:1110.4442[astro-ph.HE]

Three Nuclear forces

from Lattice QCD

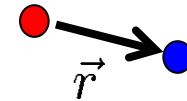


Extension from 2NF \rightarrow 3NF

In the case of 2N system...

- Calc 4pt func \rightarrow NBS amp.

$$\psi(\vec{r}) = \langle 0 | N(\vec{x} + \vec{r}; t) N(\vec{x}; t) | 2N \rangle$$



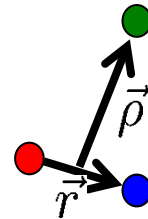
Extension to 3N system

- Calc 6pt func \rightarrow NBS amp. of 3N

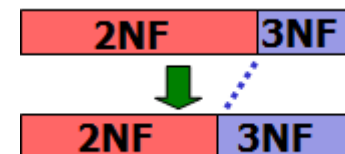
$$\psi(\vec{r}, \vec{\rho}) = \langle 0 | N(\vec{x} + \vec{r}) N(\vec{x}) N(\vec{x} + \vec{r}/2 + \vec{\rho}) | 3N \rangle$$

- Obtain 3NF through

$$(E - H_0^r - H_0^p) \psi(\vec{r}, \vec{\rho}) = \left[\underbrace{\sum_{i < j} V_{ij}(\vec{r}_{ij})}_{\substack{\uparrow \\ \text{by 2N calc}}} + V_{3NF}(\vec{r}, \vec{\rho}) \right] \psi(\vec{r}, \vec{\rho})$$



- The combination of (2NF, 3NF) \rightarrow observables
 - \rightarrow systematic determination by Lat QCD



The Challenges

- (1) S/N issue

$$S/N \sim \exp[-\mathbf{A} \times (\mathbf{m}_N - 3/2\mathbf{m}_\pi) \times \mathbf{t}]$$

- Use time-dep HAL QCD method

w/ energy-indep potential

(→ no ground state saturation is necessary)

- (2) Computational cost issue

Challenges in 3NF calculation

- Enormous computational cost for correlators

- # of Wick contraction (permutation)

$$N_{\text{perm}} = N_u! \times N_d! \sim \left[\left(\frac{3}{2} A \right)! \right]^2 \quad \text{for mass number } A$$

(← a factor of 2^A speedup by inner-baryon exchange)

- # of color / spinor contractions

$$N_{\text{loop}} = \underset{\text{(color)}}{6^A} \cdot \underset{\text{(spinor)}}{4^A} \quad (\leftarrow \text{a factor of } 2^A \text{ speedup by "half-spin" method})$$

$$N = \epsilon_{abc} (q^T C \gamma_5 q) q$$

- Total cost: $N_{\text{perm}} \times N_{\text{loop}}$

- ${}^2\text{H}$: $9 \times 144 = 1 \times 10^3$

- ${}^3\text{H}$: $360 \times 1728 = 6 \times 10^5$

- ${}^4\text{He}$: $32400 \times 20736 = 7 \times 10^8$

c.f. T.Yamazaki et al.,
PRD81(2010)111504
 $N_{\text{perm}} = 1107$ for ${}^4\text{He}$
in the isospin limit

Solution: **Unified contraction algorithm**

TD, M.Endres, arXiv:1205.0585, CPC184(2013)117

- Traditional algorithm

$$\Pi^{2N} \simeq \langle \underbrace{qqqqqq(t)}_{\text{Permutations}} \bar{q}(\xi'_1) \bar{q}(\xi'_2) \bar{q}(\xi'_3) \bar{q}(\xi'_4) \bar{q}(\xi'_5) \bar{q}(\xi'_6)(t_0) \rangle \times \underbrace{\text{Coeff}^{2N}(\xi'_1, \dots, \xi'_6)}_{\text{color } \epsilon_{abc}, \text{ spinor } (C\gamma_5), \text{ etc.}}$$

color/spinor contractions (ξ')

- New algorithm

[impose the same spacial label at source]

- Permutation applies to color/spinor indices at “Coeff”

$$\Pi^{2N} \simeq \langle qqqqqq(t) \bar{q}(\xi'_1) \bar{q}(\xi'_2) \bar{q}(\xi'_3) \bar{q}(\xi'_4) \bar{q}(\xi'_5) \bar{q}(\xi'_6)(t_0) \rangle \times \underbrace{\text{Coeff}^{2N}(\xi'_1, \dots, \xi'_6)}_{\text{Sum over color/spinor unified list}}$$

Permuted Sum

- **Permutation DONE beforehand**

- (Wick contraction and color/spinor contractions are unified)

- Significant improvement

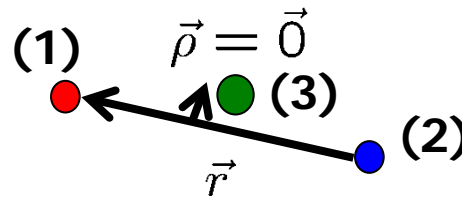
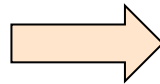
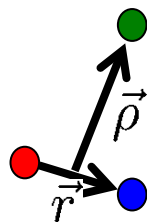


×192 for ${}^3\text{H}/{}^3\text{He}$, **×20736** for ${}^4\text{He}$, **×10¹¹** for ${}^8\text{Be}$

(x add'l. speedup)

3NF calculation in Lat QCD

- We fix the geometry of 3N (← this is not an approximation)
- We study linear setup



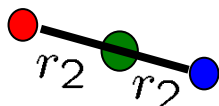
We consider
Triton channel

$$(\vec{r}_2 \equiv \vec{r}/2)$$

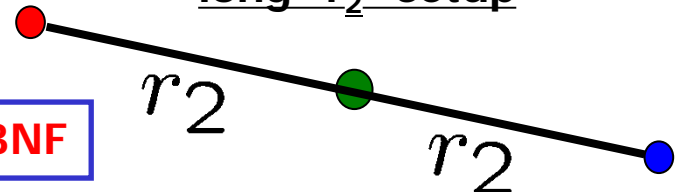
- → $L^{(1,2)\text{-pair}} = L^{\text{total}} = 0$ or 2 only
- → **Bases are only three**, labeled by $^1S_0, ^3S_1, ^3D_1$ for $(1,2)$ -pair

- **Linear setup** with various distance “ r_2 ”

short “ r_2 ” setup



long “ r_2 ” setup



Study r_2 -dependence of 3NF

Extraction of Genuine 3NF

■ **Genuine 3NF** can be extracted from **3x3 coupled channel**

- Both of parity-even 2NF and parity-odd potential required

$$\hat{H}_0 \begin{pmatrix} \psi(1S_0) \\ \psi(3S_1) \\ \psi(3D_1) \end{pmatrix} + \begin{pmatrix} V \\ (V_{2N} + V_{3NF}) \end{pmatrix} \begin{pmatrix} \psi(1S_0) \\ \psi(3S_1) \\ \psi(3D_1) \end{pmatrix} = E \begin{pmatrix} \psi(1S_0) \\ \psi(3S_1) \\ \psi(3D_1) \end{pmatrix}$$

$V_C^{I,S=1,0}, V_C^{I,S=0,1}, V_T^{I,S=0,1} : (P = \text{even})$

$V_C^{I,S=0,0}, V_C^{I,S=1,1}, V_T^{I,S=1,1} : (P = \text{odd})$

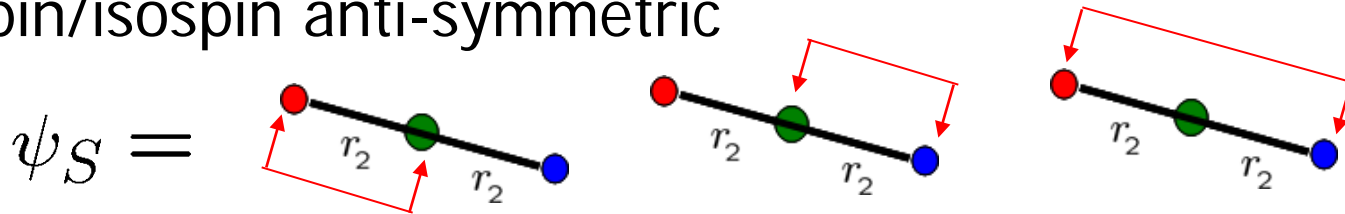
Target to be determined

■ S/N : parity-even 2NF > parity-odd 2NF in Lat QCD

- → **Desirable to extract 3NF w/ parity-even 2NF only**

Solution using “symmetric” wave function

- We can construct the wave function in which any 2N pair is spin/isospin anti-symmetric



- L=even for any 2N pair automatically guaranteed

- Bases are rotated as $|\psi_{1S_0}\rangle, |\psi_{3S_1}\rangle, |\psi_{3D_1}\rangle \rightarrow |\psi_S\rangle, |\psi_M\rangle, |\psi_{3D_1}\rangle$

$$|\psi_S\rangle = 1/\sqrt{2}(-|\psi_{1S_0}\rangle + |\psi_{3S_1}\rangle)$$

$$|\psi_M\rangle = 1/\sqrt{2}(+|\psi_{1S_0}\rangle + |\psi_{3S_1}\rangle)$$

All pair P=even

$$\hat{H}_0 \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} + \begin{pmatrix} \square & \square & \square \\ \square & V_{2N} & \square \\ \square & \square & \square \end{pmatrix} \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} + \hat{V}_{3NF} \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} = E \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix}$$

No V(P=odd)

Explicit formula for the potential matrix

- The potential matrix for the 2N part in 3x3 coupled channel in linear setup can be written as:

$$V_{2N} = \begin{pmatrix} \boxed{\begin{array}{cc} +V_C^{10}(r) + V_C^{01}(r) & +\frac{1}{2}V_C^{10}(r) - \frac{1}{2}V_C^{01}(r) \\ +\frac{1}{2}V_C^{10}(2r) + \frac{1}{2}V_C^{01}(2r) & -\frac{1}{2}V_C^{10}(2r) + \frac{1}{2}V_C^{01}(2r) \end{array}} & \begin{array}{c} -2V_T^{01}(r) \\ +2V_T^{01}(2r) \end{array} \\ \begin{array}{cc} +\frac{1}{2}V_C^{10}(r) - \frac{1}{2}V_C^{01}(r) & +\frac{3}{4}V_C^{00}(r) + \frac{1}{4}V_C^{10}(r) + \frac{1}{4}V_C^{01}(r) + \frac{3}{4}V_C^{11}(r) \\ -\frac{1}{2}V_C^{10}(2r) + \frac{1}{2}V_C^{01}(2r) & +\frac{1}{2}V_C^{10}(2r) + \frac{1}{2}V_C^{01}(2r) \end{array} & \begin{array}{c} +V_T^{01}(r) - 3V_T^{11}(r) \\ +2V_T^{01}(2r) \end{array} \\ \begin{array}{cc} -2V_T^{01}(r) & +V_T^{01}(r) - 3V_T^{11}(r) \\ +2V_T^{01}(2r) & +2V_T^{01}(2r) \end{array} & \begin{array}{c} +\frac{1}{2}V_C^{01}(r) + \frac{3}{2}V_C^{11}(r) - V_T^{01}(r) - 3V_T^{11}(r) \\ +V_C^{01}(2r) - 2V_T^{01}(2r) \end{array} \end{pmatrix} \quad (4.2)$$

$V_C^{I,S=1,0}, V_C^{I,S=0,1}, V_T^{I,S=0,1} : (P = \text{even})$

$V_C^{I,S=0,0}, V_C^{I,S=1,1}, V_T^{I,S=1,1} : (P = \text{odd})$

No $V(P=\text{odd})$

Solution using “symmetric” wave function

- We can construct the wave function in which any 2N pair is spin/isospin anti-symmetric
 - → L=even for any 2N pair automatically guaranteed
- 3x3 coupled channel is reduced to
 - one channel with only 3NF unknown
 - two channels with $V_C^{I,S=0,0}$, $V_C^{I,S=1,1}$, $V_T^{I,S=1,1}$, (3NF) unknown

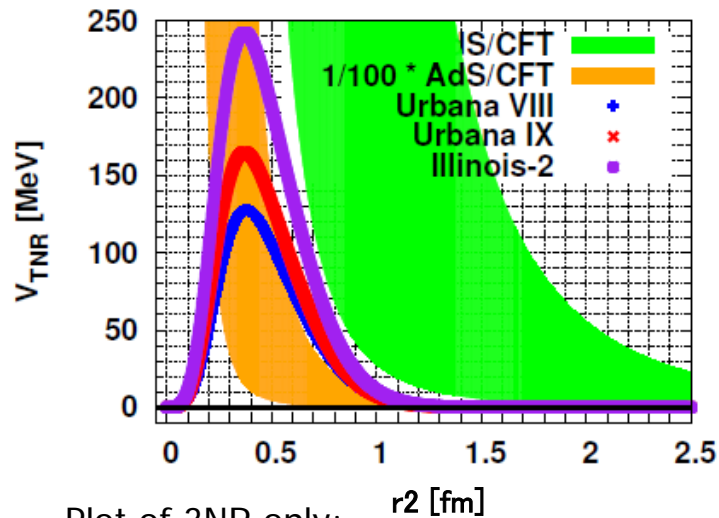
$$\begin{pmatrix} H_0 \\ \vdots \\ \vdots \end{pmatrix} \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} + \begin{pmatrix} \vdots \\ \vdots \\ V_{2N} \end{pmatrix} \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} + \begin{pmatrix} \vdots \\ \vdots \\ V_{3NF} \end{pmatrix} \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix} = E \begin{pmatrix} \psi_S \\ \psi_M \\ \psi_{3D_1} \end{pmatrix}$$

No V(P=odd)
Target to be determined

- → Even without parity-odd V, we can determine one 3NF
 - This method works for any fixed 3D-geometry other than linear

Short-Range 3NF

- We determine 3NF effectively represented by a scalar/isoscalar functional form
 - c.f. phenomenological 3NF to reproduce saturation point of nuclear matter, etc.



Plot of 3NR only:
there is cancellation from 3NA

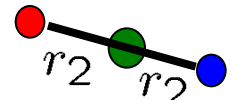
$$V_{3NF} = V_{2\pi E} + (V_{3\pi R}) + V_{3NR}$$

Urbana/Illinois

$$V_{3NR} = U_0 \sum_{cyc} T^2(r_{12})T^2(r_{13})$$

$$T(r) = \left(1 + \frac{3}{\mu r} + \frac{3}{\mu^2 r^2}\right) \frac{e^{-\mu r}}{\mu r} T_{cut}(r)$$

AdS/CFT: $V_{3NF} = +\text{const.} \cdot \frac{1}{r^4}$



K.Hashimoto, N.Iizuka
JHEP 1011 (2010) 058

Lattice QCD Calculations

Numerical Setup & Results



BG/L, BG/Q @KEK

T2K@Tsukuba



SR16000
@YITP, KEK



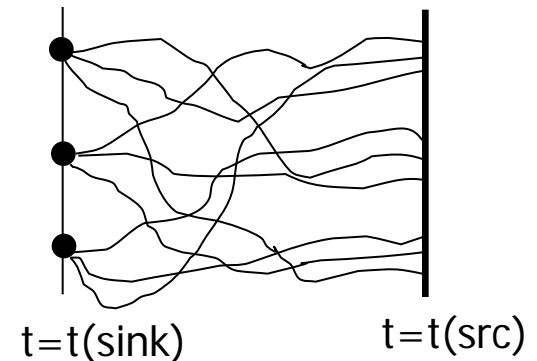
「SR16000 モデル XM1」

Lattice calculation setup

- Nf=2 clover fermion + RG improved gauge action (CP-PACS)

- 598 configs x 32 measurements
 - beta=1.95, ($a^{-1}=1.27\text{GeV}$, $a=0.156\text{fm}$)
 - $16^3 \times 32$ lattice, $L=2.5\text{fm}$
 - $M(\pi) = 1.13\text{GeV}$
 - $M(N) = 2.15\text{GeV}$ ($\kappa(\text{ud}) = 0.13750$)
 - $M(\Delta) = 2.31\text{GeV}$
- ($M\pi L=14$)

CP-PACS Coll. S. Aoki et al.,
Phys. Rev. D65 (2002) 054505



- Correlators

- **Standard nucleon op** to define the wave function / potential **at sink**

$$N = \epsilon_{abc}(q_a^T C \gamma_5 q_b) q_c$$

- **Non-rela limit op** is used to create 3N state **at source**

$$G(\vec{r}_2, t-t_0) = \sum_{\vec{x}} \langle 0 | \underbrace{N(\vec{x} + \vec{r}_2, t) N(\vec{x} - \vec{r}_2, t) N(\vec{x}, t)}_{\text{sink}} \overbrace{\overline{N N N}}_{\text{source}}(t_0) | 0 \rangle$$

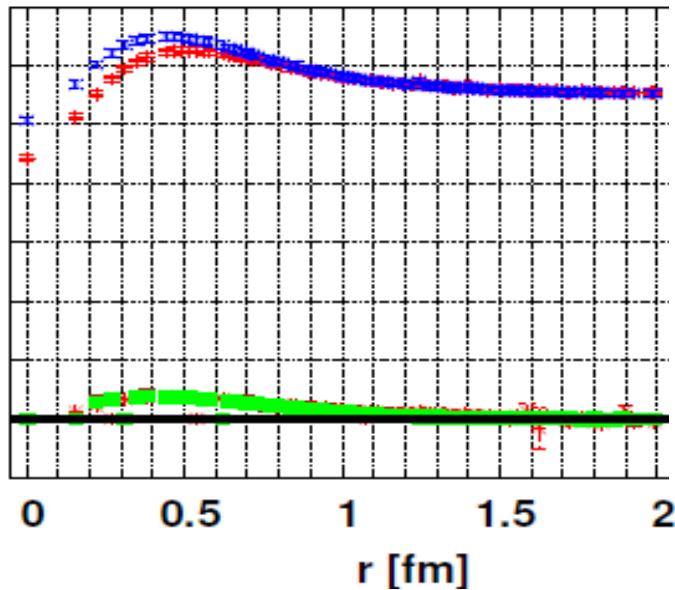
sink

source

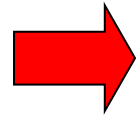
See also T. Yamazaki et al.,
PRD81(2010)111504

2NF (parity-even) from Lat QCD

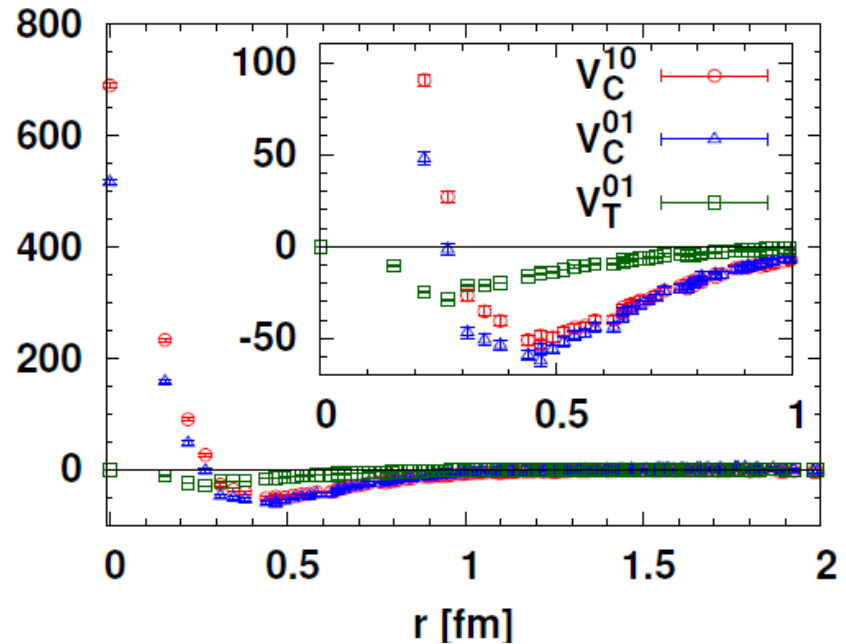
NBS Wave function



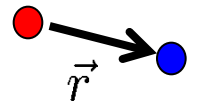
Red: 1S_0
 Blue: 3S_1
 Green: 3D_1



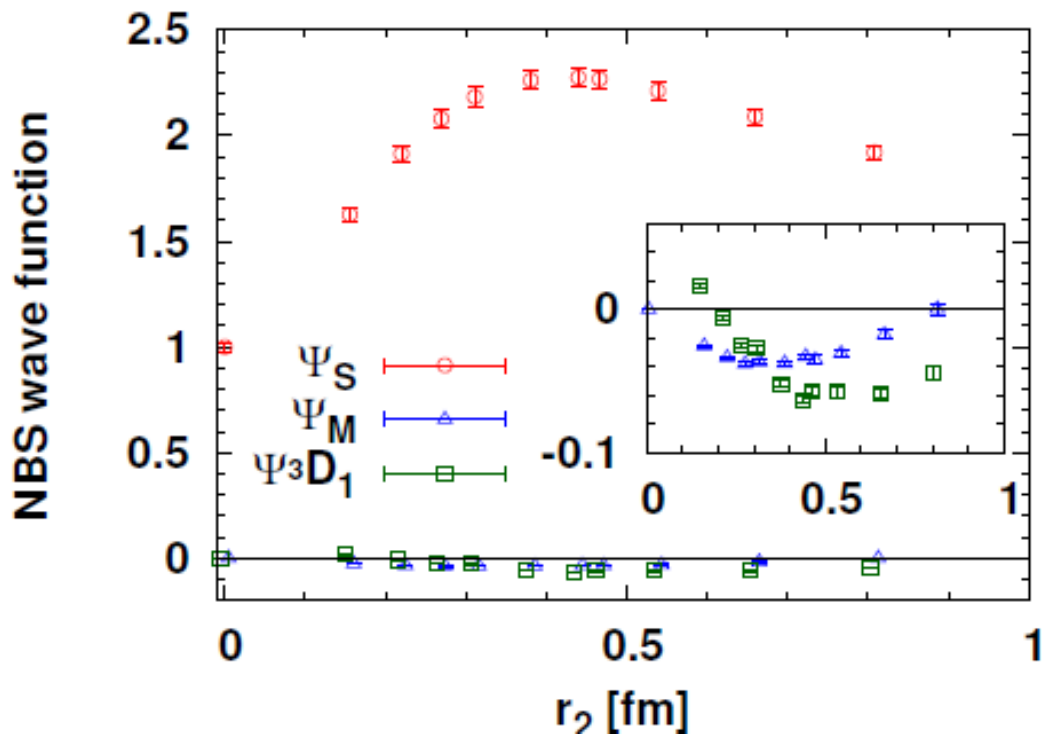
$V(r)$ [MeV]



Nf=2, CP-PACS confs, $M(\pi)=1.13\text{GeV}$



Results for wave functions

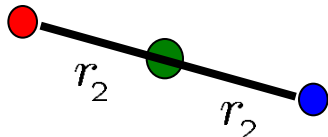


Red: Ψ_S
 Blue: Ψ_M
 Green: Ψ_{3D1}

Ψ_S overwhelms the wave function:

→ Indication of the dominance of all S-wave component, higher waves suppressed

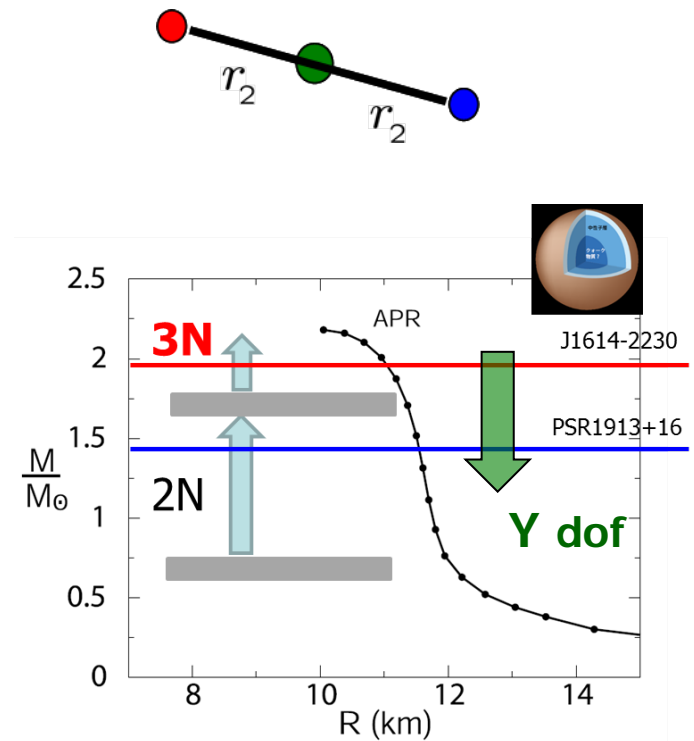
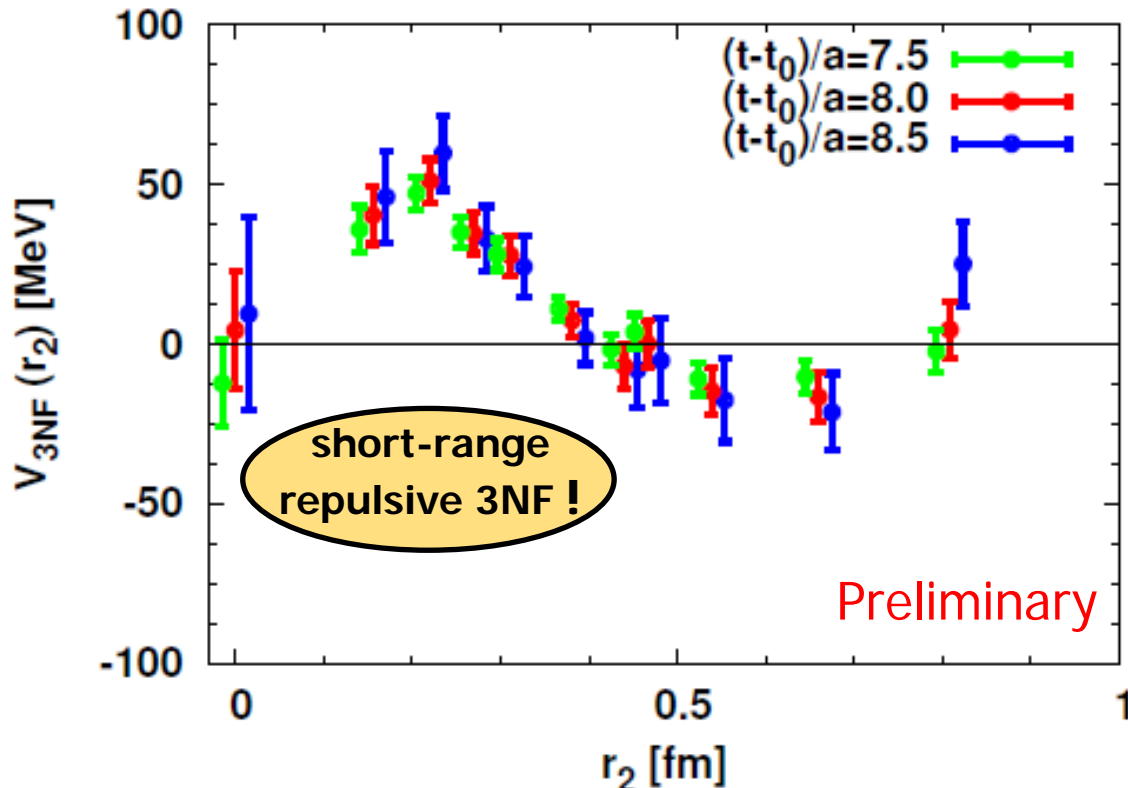
T.D. et al. (HAL QCD)
 PTP127(2012)723



3N-forces (3NF) on the lattice

T.D. et al. (HAL QCD Coll.) PTP127(2012)723

+ t-dep method updates



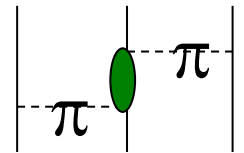
Nf=2 clover (CP-PACS), $1/a=1.27\text{GeV}$,
 $L=2.5\text{fm}$, $m_\pi=1.1\text{GeV}$, $m_N=2.1\text{GeV}$

How about other geometries ?

How about YNN, YYN, YYY ?

What is the origin of Lat 3NF ?

- 2π E-type 3NF (Fujita-Miyazawa) is unlikely
 - Strongly suppressed by $m_\pi = 1.13\text{GeV}$
- It may be attributed to quark/gluon dynamics directly
 - Recall generalized 2BF in $SU(3)_f$...

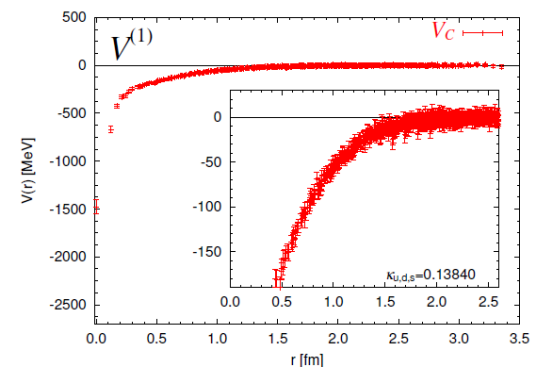
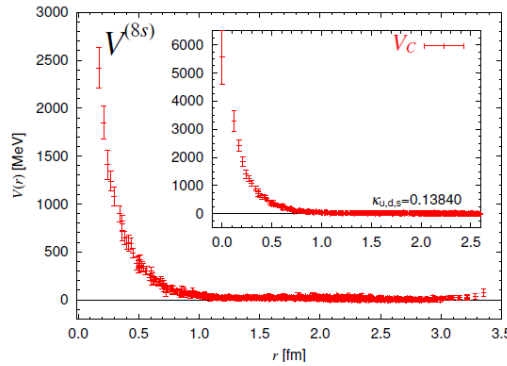
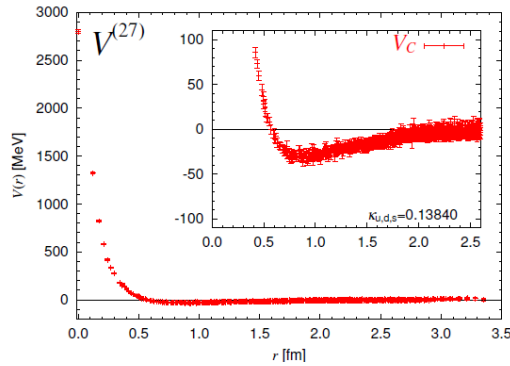


SU(3) study

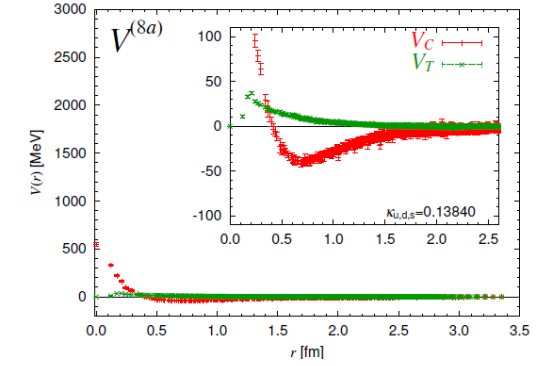
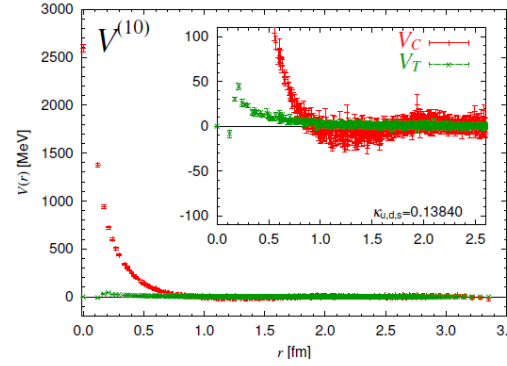
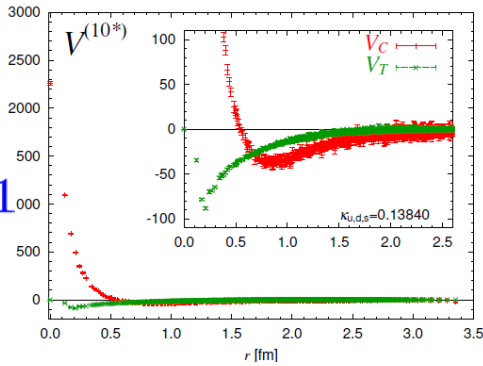
BB potentials

$a=0.12\text{fm}$, $L=3.9\text{fm}$,
 $m(\text{PS})=0.47\text{-}1.2\text{GeV}$

$1S_0$



$3S_1-3D_1$



27, 10*:
 Same as NN

8s, 10:
 strong repulsive core

1s: deep attractive pocket
8a: weak repulsive core

Repulsive core
← Pauli principle !

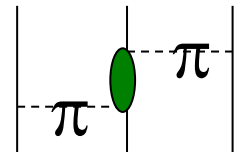
T.Inoue et al. (HAL QCD Coll.), NPA881(2012)28

Also seen in SU(2)_c, Takahashi et al., PRD82(2010)094506
 Charmonium-N, Kawanai-Sasaki, PRD82(2010)091501
 Meson-baryon, Y.Ikeda et al., arXiv:1111.2663

M.Oka et al., NPA464(1987)700

What is the origin of Lat 3NF ?

- 2π E-type 3NF (Fujita-Miyazawa) is unlikely
 - Strongly suppressed by $m_\pi = 1.13\text{GeV}$
- It may be attributed to quark/gluon dynamics directly
 - Recall generalized 2BF in $SU(3)_f$...
 - → Pauli principle works well
 - What will be **the Pauli-principle effect in 3NF** from a viewpoint of the Quark Model ?
 - c.f. OPE (pert. QCD) predicts repulsive 3NF at short distance

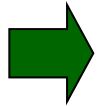


S.Aoki et al., arXiv:1112.2053

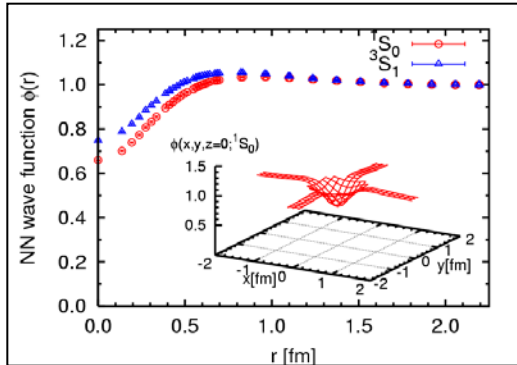
Summary

Our Approach [HAL QCD method]

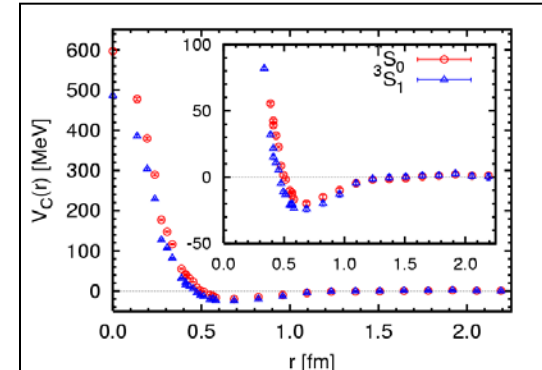
Lattice QCD



NBS wave func.



Lat Nuclear Force



$$\psi_{NBS}(\vec{r}) = \langle 0 | N(\vec{r}) N(\vec{0}) | N(\vec{k}) N(-\vec{k}), in \rangle$$

$$\simeq e^{i\delta(k)} \sin(kr + \delta(k)) / (kr)$$

(at asymptotic region)

$$(k^2/m_N - H_0) \psi(\vec{r}) = \int d\vec{r}' U(\vec{r}, \vec{r}') \psi(\vec{r}')$$

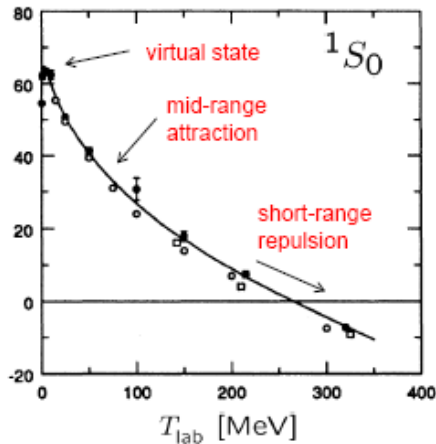
Lat potential is faithful to phase shift by construction

Analog to ...

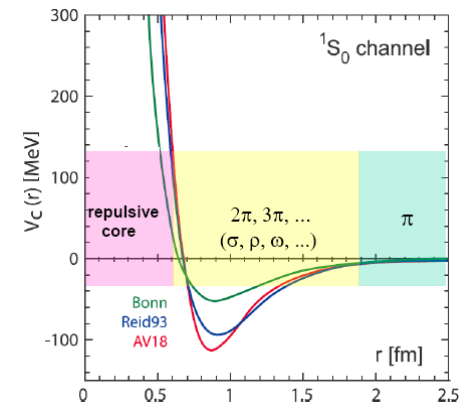
Scattering Exp.



Phase shifts

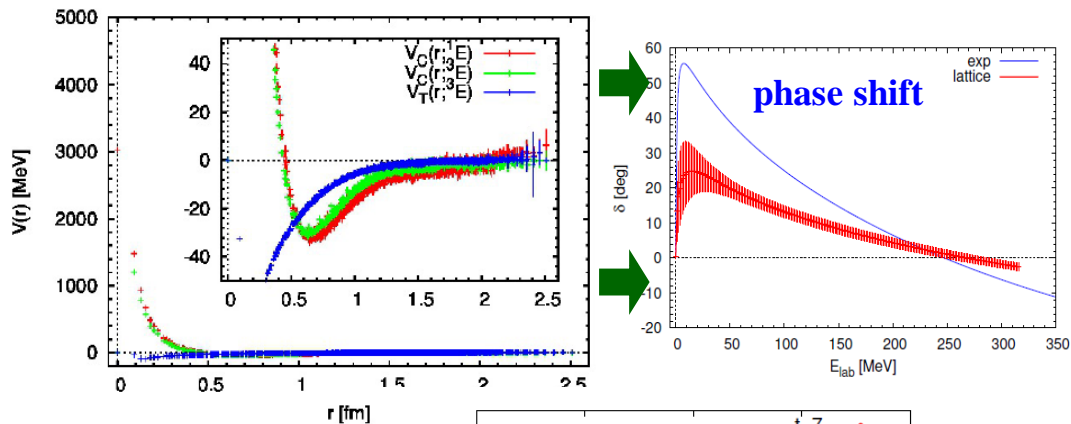


Phen. Potential

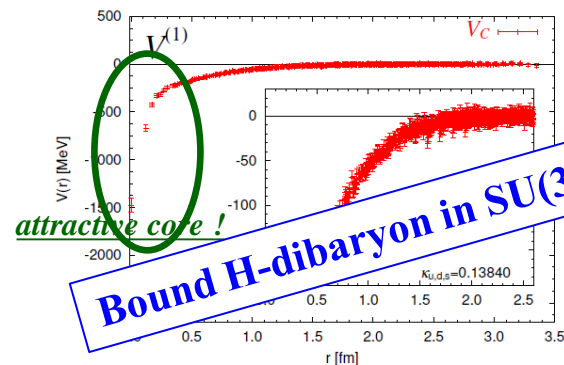


Research Highlight

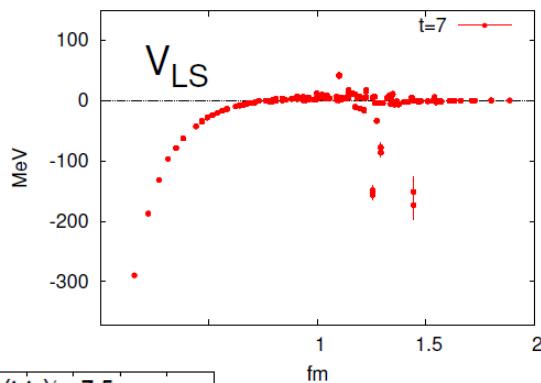
NN forces (central & tensor)



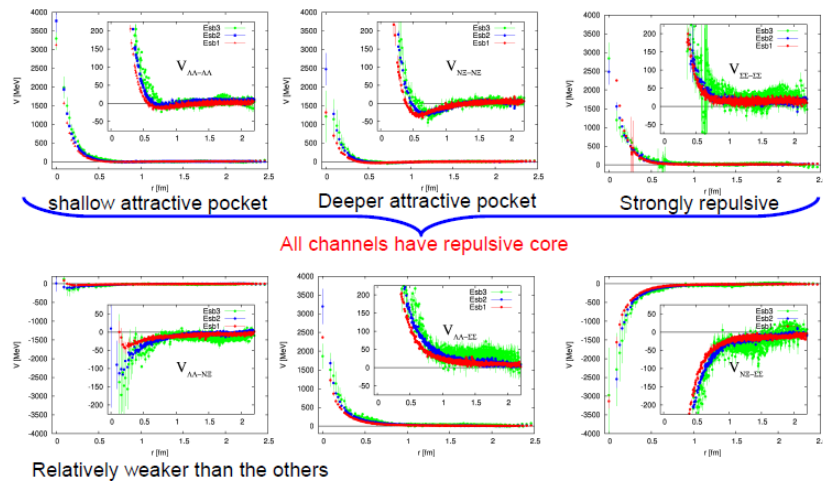
Hyperon Forces



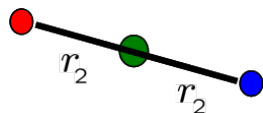
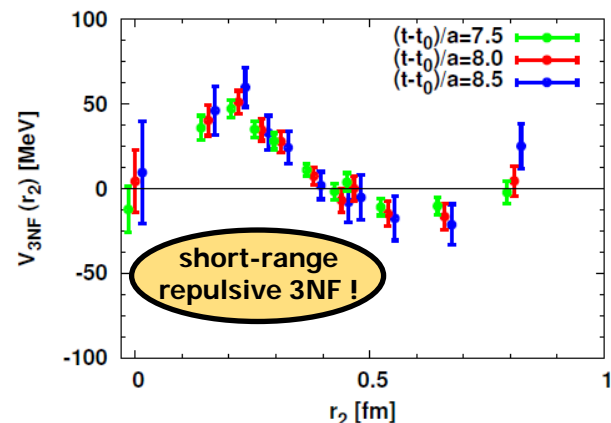
P-odd forces, LS force



Coupled channel beyond SU(3)



Three-Nucleon Forces



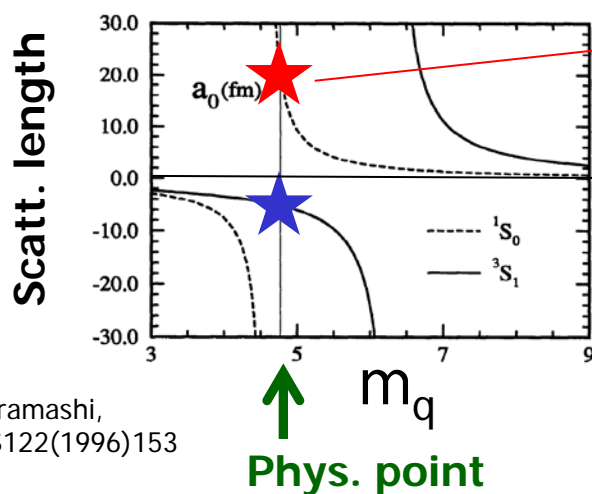
Future Prospects in Lattice Nuclear Forces

- *What is most important next ?*

Towards realistic potential on the Lattice

• **Physical mass point**, Infinite V limit, continuum limit

– Physical m_π crucial for OPEP, chiral extrapolation won't work



$a_0(^1S_0) \sim 20\text{fm}$

“Unitary Region”



We are here

→ m_q

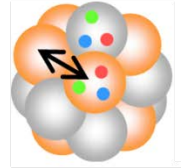
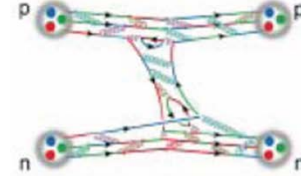
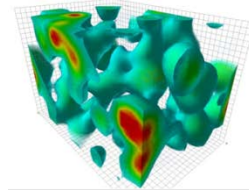
Y. Kuramashi, PTSP122(1996)153

– Gauge confs generation at $m_\pi = 140\text{MeV}$, $L \sim 10\text{fm}$ on the **K computer**



K computer

Summary and Prospects



- **Hadron Interactions** by 1st principle Lat calc
 - Bridging different worlds:
Particle Physics / **Nuclear Physics** / **Astrophysics**
- Lattice QCD results for **NN**, **YN/YY**, **NNN**
 - **Intriguing physics** even at heavy quark masses
- **Next major step: physical quark mass point !**
 - Breakthroughs in **S/N issue** & **Comput. cost issue**

**YOUR new idea
Welcome !**

Thermodynamic limit & continuum limit

- ➔ **Realistic hadron interactions**
- ➔ **Nuclear Physics on the Lattice !**